

#### DART\_LAB Tutorial Section 1: Ensemble Data Assimilation Concepts in 1D







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#### What is Data Assimilation?

#### Observations combined with a Model forecast ...



An observation has a value (\*),



An observation has a value (\*),



and an error distribution (red curve) that is associated with the instrument.

Thermometer outside measures 1°C.



Instrument builder says thermometer is unbiased with +/- 0.8° C gaussian error.

Thermometer outside measures 1°C.



The red plot is  $P(T \mid T_0)$ ;

probability of temperature given that  $T_o$  was observed.

We also have a prior estimate of temperature.



The green curve is P(T | C);

probability of temperature given all available prior information C.

Prior information *C* can include:

- 1. Observations of things besides T;
- 2. Model forecast made using observations at earlier times;
- 3. *a priori* physical constraints (T > -273.15°C);
- 4. Climatological constraints ( $-30^{\circ}C < T < 40^{\circ}C$ ).

**Likelihood**: Probability that  $T_o$  is observed if T is true value and given prior information C. Theorem:  $P(T | T_o, C) = \frac{P(T_o | T, C)P(T | C)}{P(T_o | C)}$ 

**Posterior**: Probability of *T* given observations and Prior. Also called **update** or **analysis**.

**Rewrite Bayes as:** 

$$\frac{P(T_o \mid T, C)P(T \mid C)}{P(T_o \mid C)} = \frac{P(T_o \mid T, C)P(T \mid C)}{\int P(T_o \mid x)P(x \mid C)dx}$$
$$= \frac{P(T_o \mid T, C)P(T \mid C)}{normalization}$$

Denominator normalizes so Posterior is PDF.

$$P(T | T_0, C) = \frac{P(T_0 | T, C)P(T | C)}{normalization}$$











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## **Consistent Color Scheme Throughout Tutorial**

- Green = Prior
- Red = Observation
- Blue = Posterior

Black = Truth

(truth available only for 'perfect model' examples)

$$P(T | T_0, C) = \frac{P(T_0 | T, C)P(T | C)}{normalization}$$

Generally no analytic solution for Posterior.



DART\_LAB Section 1: 17 of 69

$$P(T | T_0, C) = \frac{P(T_0 | T, C)P(T | C)}{normalization}$$

Gaussian Prior and Likelihood -> Gaussian Posterior



For Gaussian prior and likelihood...

- Prior  $P(T | C) = Normal(T_p, \sigma_p)$
- Likelihood  $P(T_o | T, C) = Normal(T_o, \sigma_o)$
- Then, Posterior

$$P(T \mid T_o, C) = Normal(T_u, \sigma_u)$$

With

$$T_u = \sigma_u^2 \left[ \sigma_p^{-2} T_p + \sigma_o^{-2} T_o \right]$$

 $\sigma_u = \sqrt{\left(\sigma_p^{-2} + \sigma_o^{-2}\right)^{-1}}$ 

## Matlab Hands-on: gaussian\_product



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**Purpose**: Explore the gaussian posterior that results from taking the product of a gaussian prior and a gaussian likelihood.



## Matlab Hands-on: gaussian\_product

#### Explore!

- Change the mean value of the prior and the observation.
- Change the standard deviation of the prior.
- What is always true for the mean of the posterior?
- What is always true for the standard deviation of the posterior?

- 1. Suppose we have a linear forecast model L
  - A. If temperature at time  $t_1 = T_{1,}$  then the temperature at  $t_2 = t_1 + \Delta t$  is  $T_2 = L(T_1)$
  - B. Example:  $T_2 = T_1 + \Delta t T_1$

- 1. Suppose we have a linear forecast model L.
  - A. If temperature at time  $t_1 = T_{1_1}$  then the temperature at  $t_2 = t_1 + \Delta t$  is  $T_2 = L(T_1)$
  - B. Example:  $T_2 = T_1 + \Delta t T_1$
- 2. If posterior estimate at time  $t_1$  is *Normal*( $T_{u,1}, \sigma_{u,1}$ ) then the prior at  $t_2$  is *Normal*( $T_{p,2}, \sigma_{p,2}$ ).

$$T_{p,2} = T_{u,1}, + \Delta t T_{u,1}$$
$$\sigma_{p,2} = (\Delta t + 1) \sigma_{u,1}$$

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- 3. Given an observation at  $t_2$  with distribution *Normal*( $t_0, \sigma_0$ ) the likelihood is also *Normal*( $t_0, \sigma_0$ ).

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- 3. Given an observation at  $t_2$  with distribution *Normal*( $t_0, \sigma_0$ ) the likelihood is also *Normal*( $t_0, \sigma_0$ ).
- 4. The posterior at  $t_2$  is *Normal*( $T_{u,2}, \sigma_{u,2}$ ) where  $T_{u,2}$  and  $\sigma_{u,2}$  come from page 19.

#### A One-Dimensional Ensemble Kalman Filter

Represent a prior pdf by a sample (ensemble) of N values:



## A One-Dimensional Ensemble Kalman Filter

Represent a prior pdf by a sample (ensemble) of N values:



DART\_LAB Section 1: 28 of 69

If posterior ensemble at time  $t_1$  is  $T_{1,n}$ , n = 1, ..., N



If posterior ensemble at time  $t_1$  is  $T_{1,n}$ , n = 1, ..., Nadvance each member to time  $t_2$  with model,  $T_{2,n} = L(T_{1,n})$  n = 1, ..., N.



Same as advancing continuous pdf at time  $t_1 \dots$ 



Same as advancing continuous pdf at time  $t_1 \dots$  to time  $t_2$  with model L.







Fit a Gaussian to the sample.



Get the observation likelihood.



Compute the continuous posterior PDF.



Use a deterministic algorithm to 'adjust' the ensemble.



First, 'shift' the ensemble to have the exact mean of the posterior.



First, 'shift' the ensemble to have the exact mean of the posterior. Second, linearly contract to have the exact variance of the posterior. Sample statistics are identical to Kalman filter.

## Matlab Hands-On: oned\_ensemble



## Matlab Hands-On: oned\_ensemble

#### Purpose: Explore how ensemble filters update a prior ensemble.



## Matlab Hands-On: oned\_ensemble

#### **Explorations:**

- 1. Keep your ensembles small, less than 10, for easy viewing.
- Create a nearly uniformly spaced ensemble.
  Examine the update.
- 3. What happens with an ensemble that is confined to one side of the likelihood?
- 4. What happens with a bimodal ensemble (two clusters of members on either side)?
- 5. What happens with a single outlier in the ensemble?



Purpose:

- Explore behavior of a complete 1-D ensemble filter for a linear system.
- Look at the behavior of different ensemble sizes.



The 'truth' is always 0.

Observation noise is a draw from N(0,1).

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A time series of the assimilation. Line segments show forecast evolution. Most recent prior, observation, and posterior are same as in upper right window but plotted with a vertical axis.

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Time series of error and spread. Sawtooth pattern because prior and posterior values are shown for each time.

Purpose:

- Explore behavior of a complete 1-D ensemble filter for a linear system.
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Explorations:

- Step through a sequence of advances and assimilations with the top button. Watch the evolution of the ensemble, the error and spread.
- 2. How does a larger ensemble size ( < 10 is easiest to see) act?
- Compare the error and spread for different ensemble sizes.
- Note the time behavior of the error and spread.
- 3. Let the model run freely using the second button.

Draw 5 values from a real-valued distribution. Call the first 4 'ensemble members'.



These 4 'ensemble members' partition the real line into 5 bins.



Call the 5th draw the 'truth'. 1/5 chance that this is in any given bin.











Rank histograms for good ensembles should be uniform (caveat sampling noise). Want truth to look like random draw from ensemble.



A biased ensemble leads to skewed histograms.



An ensemble with too little spread gives a u-shape. This is the most common behavior for geophysics.



An ensemble with too much spread is peaked in the center.



# Matlab Hands-On: **oned\_model** Understanding the Rank Histogram

Truth is always 0. For this time, truth is between 1<sup>st</sup> and 2<sup>nd</sup> ensemble members; that's the second bin.

Prior (left) and Posterior (right) rank histograms. Yellow is for current time.



# Matlab Hands-On: oned\_model Understanding the Rank Histogram

#### Explorations:

- Step through a sequence of advances and assimilations with the top button. Watch the evolution of the rank histogram bins.
- 2. Add some model bias (less than 1 to start) and see how the filter responds.
- 3. Add some nonlinearity ( a< 1 ) to the model. How do the different filters respond?
- 4. Can you break the filter (find setting so that the ensemble moves away from zero) with the options explored so far?

Observations + physical system \_\_\_\_\_\_ 'true' distribution.





Could correct error if we knew what it was.

With large models, can't know error precisely.

Taking no action can cause observations to be ignored.



Naive solution: increase the spread in the prior.

Give more weight to the observation, less to the prior.



# Matlab Hands-On: **oned\_ensemble** exploring prior inflation



# Matlab Hands-On: exploring prior inflation with oned\_ensemble

Explorations:

- See how increasing inflation (> 1) changes the posterior mean and standard deviation.
- Look at priors that are not shifted but have small spread compared to the observation error distribution.
- Look at priors that are shifted from the observation.

# Matlab Hands-On: **oned\_model** using inflation to deal with systematic error



- 1. Add some model bias to simulate systematic error.
  - Run an assimilation and observe the error, spread, and rank histograms.
- 3. Add some inflation
  - (try starting with 1.5) and observe how behavior changes.
- 4. What happens with too much inflation?

Note: The spread is increased by the square root of the inflation.

# Matlab Hands-On: oned\_model using inflation to deal with systematic error

#### **Explorations:**

- Try a variety of model bias and inflation settings.
- Try using inflation with a nonlinear model.

