

DART Tutorial Section 6: Other Updates for an Observed Variable







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How can we take product of sample with continuous likelihood?



$$p(A \mid BC) = \frac{p(B \mid AC)p(A \mid C)}{p(B \mid C)} = \frac{p(B \mid AC)p(A \mid C)}{\int p(B \mid x)p(x \mid C)dx}$$

Observation likelihood usually continuous (nearly always Gaussian).



If Obs. Likelihood isn't Gaussian, can generalize methods below. For instance, can fit set of Gaussian kernels to obs. likelihood.

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Product of prior Gaussian fit and Obs. likelihood is Gaussian.



Computing continuous posterior is simple. BUT, need to have a SAMPLE of this PDF.

There are many ways to do this.



Exact properties of different methods may be unclear. Trial and error still best way to see how they perform. Will interact with properties of prediction models, etc.

Just draw a random sample (*filter_kind=5* in &assim_tools_nml).



NOTE: When trying filter_kinds other than 1, sort_obs_inc in &assim_tools_nml should be .true. (see section 10).

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Can 'play games' with this sample to improve (modify) its properties. Example: Adjust the mean of the sample to be exact.

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Can 'play games' with this sample to improve (modify) its properties.

Example: Adjust the mean of the sample to be exact. Can also adjust the variance to be exact.

Just draw a random sample (*filter_kind=5* in &assim_tools_nml).



Might also want to eliminate rare extreme outliers.

NOTE: Properties of these adjusted samples can be quite different. How these properties interact with the rest of the assimilation is an open question.

Construct a 'deterministic' sample with certain features.



For instance: Sample could have exact mean and variance.

This is insufficient to constrain ensemble, need other constraints.

Construct a 'deterministic' sample with certain features (*filter_kind*=6 in &assim_tools_nml; manually adjust kurtosis).



Example: Exact sample mean and variance.

Sample kurtosis is 3 (expected value for Gaussian in large sample limit) (Constructed by starting uniformly-spaced and adjusting quadratically).

Construct a 'deterministic' sample with certain features (*filter_kind*=6 in &assim_tools_nml; manually adjust kurtosis).



Example: Exact sample mean and variance.

Sample kurtosis 2: less extreme outliers, less dense near mean. Avoiding outliers might be nice in certain applications. Sampling heavily near mean might be nice.

First two methods depend only on mean and variance of prior sample.



Example: Suppose prior sample is (significantly) bimodal?

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Example: Suppose prior sample is (significantly) bimodal?

Might want to retain additional information from prior. Recall that Ensemble Adjustment Filter tried to do this (Section 1).

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



'Classical' Monte Carlo Algorithm for Data Assimilation. Warning: earliest references (Evensen 1994) have suboptimal algorithm (more in a minute).

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Again, fit a Gaussian to the sample. Are there ways to do this without computing prior sample stats?

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Again, fit a Gaussian to the sample.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Generate a random draw from the observation likelihood. Associate it with the first sample of the prior ensemble.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Proceed to associate a random draw from obs. with each prior sample. This has been called 'perturbed' observations. Algorithm sometimes called 'perturbed obs.' ensemble Kalman filter.

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Proceed to associate a random draw from obs. with each prior sample. Earliest publications associated mean of obs. likelihood with each prior. This resulted in insufficient variance in posterior.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Have sample of joint prior distribution for observation and prior MEAN.Adjusting the mean of obs. sample to be exact improves performance.Adjusting the variance may further improve performance.Outliers are potential problem, but can be removed.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



For each prior mean/obs. pair, find mean of posterior PDF.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Prior sample standard deviation still measures uncertainty of prior mean estimate.

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Prior sample standard deviation still measures uncertainty of prior mean estimate. Obs. likelihood standard deviation measures uncertainty of obs. estimate.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Take product of the prior/obs distributions for first sample. This is the standard Gaussian product.

Ensemble Kalman Filter (EnKF) (*filter_kind*=2 in &assim_tools_nml).



Mean of product is random sample of posterior. Product of random samples is random sample of product.

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Posterior sample maintains much of prior sample structure. (This is more apparent for larger ensemble sizes.) Posterior sample mean and variance converge to 'exact' for large samples.

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



Can retain more correct information about non-Gaussian priors. Can also be used for obs. likelihood term in product (not shown here).

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



Usually, kernel widths are a function of the sample variance. Almost avoids using prior sample variance.

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The estimate of the prior is the normalized sum of all kernels.

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



Apply distributive law to take product. Product of the sum is the sum of the products.

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



Compute product of first kernel with Obs. Likelihood.

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



But, can no longer ignore the weight term for product of Gaussians. Kernels with mean further from observation get less weight.

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



Continue to take products for each kernel in turn. More distant kernels have smaller impact on posterior.

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Final posterior is weight-normalized sum of kernel products.

Posterior is somewhat different than for ensemble adjustment or ensemble Kalman filter (much less density in left lobe.)

Ensemble Kernel Filter (EKF) (*filter_kind*=3 in &assim_tools_nml).



Forming sample of the posterior can be problematic. Random sample is simple.

Deterministic sampling is much more tricky here (few results available).

Rank Histogram Filter (*filter_kind=8* in &assim_tools_nml).



Goal: Want to handle non-Gaussian priors or observation likelihoods. Low information content obs. must yield small increments. Must perform well for Gaussian priors. Must be computationally efficient.

Anderson, J. L., 2010: A Non-Gaussian Ensemble Filter Update for Data Assimilation. *Mon. Wea. Rev.*, **139**, 4186-4198. doi: 10.1175/2010MWR3253.1 DART Tutorial Section 6: Slide 56

Rank Histogram Filter (*filter_kind=8* in &assim_tools_nml).



- Place (ens_size + 1)⁻¹ mass between adjacent ensemble members.
- Reminiscent of rank histogram evaluation method.

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- Partial Gaussian kernels on tails, N(*tail_mean, ens_sd*).
- tail_mean selected so that (ens_size + 1)⁻¹ mass is in tail.

Rank Histogram Filter (*filter_kind*=8 in &assim_tools_nml).



Step 2: Use likelihood to compute weight for each ensemble member.

- Analogous to classical particle filter.
- Can be extended to non-Gaussian obs. likelihoods.

Rank Histogram Filter (*filter_kind*=8 in &assim_tools_nml).



Step 2: Use likelihood to compute weight for each ensemble member.

• Can approximate interior likelihood with linear fit; for efficiency.

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Step 3: Compute continuous posterior distribution.

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Step 3: Compute continuous posterior distribution.

- Product of prior Gaussian kernel with likelihood for tails.
- Easy for Gaussian likelihood.
- More quadrature if non-Gaussian likelihood.

Rank Histogram Filter (*filter_kind=8* in &assim_tools_nml).



Step 4: Compute updated ensemble members:

- (ens_size +1)⁻¹ of posterior mass between each ensemble pair.
- (ens_size +1)⁻¹ in each tail.
- Uninformative observation has no impact.

Rank Histogram Filter (*filter_kind*=8 in &assim_tools_nml).



Compare to standard Ensemble Adjustment Filter (EAKF). Nearly Gaussian case, differences are small.

Rank Histogram Filter (*filter_kind=8* in &assim_tools_nml).



Rank Histogram gets rid of outlier that is clearly inconsistent with obs. EAKF can't get rid of outlier.

Large prior variance from outlier causes EAKF to shift all members too much towards observation.

Rank Histogram Filter (*filter_kind=8* in &assim_tools_nml).



Rank Histogram handles multimodal prior and compelling observation. EAKF still bimodal; left mode is inconsistent with everything. Lorenz 63 can have priors like this.
Rank Histogram Filter (*filter_kind*=8 in &assim_tools_nml).



Convective-scale models (and land models) have analogous behavior. Convection may fire at 'random' locations.

Subset of ensembles will be in right place, rest in wrong place.

Want to aggressively eliminate convection in wrong place.

Particle filter methods:

These are 'classical' ensemble methods from statistical literature.

Size of ensembles required scales hyper-exponentially with model size.

Ensembles > 1000 required for models with > 4 degrees of freedom.

This rules out naive application to any meaningful atmospheric model.

Particle Filter (*filter_kind*=4 in &assim_tools_nml).

Can use particle filters in a few dimensions.

DART provides a one-dimensional particle filter.

Independent particle filter is used for updating each observation.

PROBLEM: Inconsistency between updates for different observations.

This can probably be made to work in some clever way!

Other novel filter algorithms available on request:

- 1. Quadratic filter by Dan Hodyss, NRL
- 2. Localized particle filter by Jon Poterjoy, NOAA

Contact the DART development team for more information, dart@ucar.edu.

Using DART Diagnostics

- What happens when these different methods are used in Lorenz 63?
- Are they significantly different?
- Do some work better for different observation sets?
- Can kernel filter deal better with distinct bimodality of Lorenz 63? With proper resampling, this should be the case. Somebody clever could probably make this work.

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